

Perturbative Quantum Gravity

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Review of some of modern developments in the subject

(designed for a mathematical audience)

Why mathematicians are an appropriate audience?

- (Perturbative) QG is a mess if approached in a brute force way
- Simplicity every time one gets through the mess

Everything this subject needs is some level of
mathematical sophistication?

Why is it relevant to study this subject now?

- Abandoned in the 80's in favour of other approaches to quantum gravity, notably string theory
- Supersymmetry (needed by superstring) was widely expected to be discovered by the new generation of accelerators
- No SUSY was discovered at LHC
- Possible that the Standard Model of elementary particles + Gravity is everything there is up to the Planck scale

It is time to re-evaluate the status of the field theoretic version of quantum gravity

Outline of the talk

- Review of gravitational perturbation theory and quantization
- New beautiful results about “gravitons” in the last 10 years

“On-shell methods”

Simplicity points to some
underlying structure yet to be
discovered

- New gauge-theoretic formulation of gravity

Simplifies perturbation theory

Suggests a fresh look at old problems

Classical General Relativity

Spacetime dimension $D=4$

Consider “pure gravity”

dynamical theory of a spacetime metric

World without
material sources

$$S_{\text{EH}}[g] = \frac{1}{16\pi G} \int \sqrt{-\det(g)} (R - 2\Lambda)$$

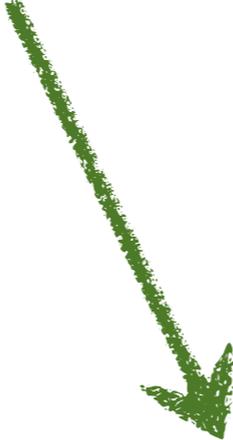
Einstein-Hilbert action

R - scalar curvature

Λ - cosmological constant

G - Newton's constant

Einstein metrics are critical
points of EH functional


$$R_{\mu\nu} = \Lambda g_{\mu\nu}$$

Highly non-linear second order PDE on $g_{\mu\nu}$

Perturbation theory

Set $\Lambda = 0$ for now

Minkowski metric

$$\eta_{\mu\nu} = \text{diag}(-1, 1, 1, 1)$$

Physically well-motivated
because Λ is small

is a solution of field equations

$$\kappa^2 := 32\pi G$$

Expand $g_{\mu\nu} = \eta_{\mu\nu} + \kappa h_{\mu\nu}$

set $\kappa = 1$ from now on

$$\mathcal{L}^{(2)} = -\frac{1}{2}(\partial_\rho h_{\mu\nu})^2 + \frac{1}{2}(\partial_\mu h)^2 + (\partial^\nu h_{\mu\nu})^2 + h\partial^\mu\partial^\nu h_{\mu\nu}$$

where $h := \eta^{\mu\nu} h_{\mu\nu}$

Gravitons

Define $\bar{h}_{\mu\nu} := h_{\mu\nu} - \frac{1}{2}h\eta_{\mu\nu}$

Linearized field equations

where

$$\square := \partial^\mu \partial_\mu$$

$$\square \bar{h}_{\mu\nu} + \eta_{\mu\nu} \partial^\mu \partial^\nu \bar{h}_{\mu\nu} - 2\partial_{(\mu} \partial^\rho \bar{h}_{\nu)\rho} = 0$$

Everything is invariant under diffeomorphisms

$$\delta h_{\mu\nu} = 2\partial_{(\mu} \xi_{\nu)}$$

Can “fix the gauge”

$$\partial^\nu \bar{h}_{\mu\nu} = 0$$

Field equations become

$$\square \bar{h}_{\mu\nu} = 0$$

solutions-
plane waves-
gravitons

Mode expansion

only two polarizations propagate

amplitudes of two
different polarizations

2+2 numbers per spatial
point as initial data

$$h_{\mu\nu}^k = a_k^+ \epsilon_{\mu\nu}^+(k) e^{ikx} + a_k^- \epsilon_{\mu\nu}^-(k) e^{ikx}$$

two polarization tensors

Polarization tensors satisfy

$$\begin{aligned} (\epsilon_{\mu\nu}^\pm(k))^2 &= 0 \\ \epsilon_{\mu\nu}^+(k) \epsilon^{-\mu\nu}(k) &= 1 \end{aligned}$$

$\left(\begin{array}{c} \text{Positive} \\ \text{Negative} \end{array} \right)$ helicity \equiv only $\left(\begin{array}{c} \text{Self-dual} \\ \text{Anti-self-dual} \end{array} \right)$ part of Weyl is non-vanishing

Einstein gravity perturbatively: Nasty mess...

Expansion around an arbitrary background $g_{\mu\nu}$

quadratic order (together with the gauge-fixing term)

$$L_{g.f.} = -\sqrt{-g} \left(h^{\mu\nu}{}_{;\nu} - \frac{1}{2} h_{\nu}{}^{\nu;\mu} \right) \left(h^{\rho}{}_{\mu;\rho} - \frac{1}{2} h^{\rho}{}_{\rho;\mu} \right)$$

$$L_2 = \sqrt{-g} \left\{ -\frac{1}{2} h^{\alpha\beta}{}_{;\gamma} h_{\alpha\beta}{}^{;\gamma} + \frac{1}{4} h^{\alpha}{}_{\alpha;\gamma} h_{\beta}{}^{\beta;\gamma} + h_{\alpha\beta} h_{\gamma\delta} R^{\alpha\gamma\beta\delta} - h_{\alpha\beta} h^{\beta}{}_{\gamma} R^{\delta\alpha\gamma}{}_{\delta} \right. \\ \left. + h^{\alpha}{}_{\alpha} h_{\beta\gamma} R^{\beta\gamma} - \frac{1}{2} h_{\alpha\beta} h^{\alpha\beta} R + \frac{1}{4} h^{\alpha}{}_{\alpha} h^{\beta}{}_{\beta} R \right\}.$$

from Goroff-Sagnotti
"2-loop" paper

cubic order

$$L_3 = \sqrt{-g} \left\{ -\frac{1}{2} h^{\alpha\beta} h^{\gamma\delta}{}_{;\alpha} h_{\gamma\delta;\beta} + 2h^{\alpha\beta} h^{\gamma\delta}{}_{;\alpha} h_{\beta\gamma;\delta} - h^{\alpha\beta} h^{\gamma}{}_{\gamma;\alpha} h^{\delta}{}_{\delta;\beta} - \frac{1}{2} h^{\alpha}{}_{\alpha} h^{\beta\gamma;\delta} h_{\beta\delta;\gamma} \right. \\ \left. + \frac{1}{4} h^{\alpha}{}_{\alpha} h^{\beta\gamma;\delta} h_{\beta\gamma;\delta} - h^{\alpha\beta} h^{\gamma}{}_{\gamma;\delta} h^{\delta}{}_{\alpha;\beta} + \frac{1}{2} h^{\alpha\beta} h^{\gamma}{}_{\gamma;\alpha} h^{\delta}{}_{\delta;\beta} - h^{\alpha\beta} h_{\alpha\beta;\gamma} h^{\gamma\delta}{}_{;\delta} \right. \\ \left. + \frac{1}{2} h^{\alpha}{}_{\alpha} h^{\beta}{}_{\beta;\gamma} h^{\gamma\delta}{}_{;\delta} + h^{\alpha\beta} h_{\alpha\beta;\gamma} h^{\delta}{}_{\delta;\gamma} + \frac{1}{4} h^{\alpha}{}_{\alpha} h^{\beta}{}_{\beta;\gamma} h^{\delta}{}_{\delta;\gamma} - h^{\alpha\beta} h^{\gamma}{}_{\alpha;\delta} h_{\beta\gamma}{}^{;\delta} \right. \\ \left. + h^{\alpha\beta} h^{\gamma}{}_{\alpha;\delta} h^{\delta}{}_{\beta;\gamma} + R_{\alpha\beta} (2h^{\alpha\gamma} h_{\gamma\delta} h^{\beta\delta} - h^{\gamma}{}_{\gamma} h^{\alpha\delta} h^{\beta}{}_{\delta} - \frac{1}{2} h^{\alpha\beta} h^{\gamma\delta} h_{\gamma\delta} \right. \\ \left. + \frac{1}{4} h^{\alpha\beta} h^{\gamma}{}_{\gamma} h^{\delta}{}_{\delta}) + R \left(-\frac{1}{3} h^{\alpha\beta} h_{\beta\gamma} h^{\gamma}{}_{\alpha} + \frac{1}{4} h^{\alpha}{}_{\alpha} h^{\beta\gamma} h_{\beta\gamma} - \frac{1}{24} h^{\alpha}{}_{\alpha} h^{\beta}{}_{\beta} h^{\gamma}{}_{\gamma} \right) \right\}$$

even in flat space, the corresponding vertex has about 100 terms!

quartic order

$$\begin{aligned}
 L_4 = \sqrt{-g} \left\{ & (h^\alpha_\alpha h^\beta_\beta - 2h^{\alpha\beta} h_{\alpha\beta}) \left(\frac{1}{16} h^{\gamma\delta;\sigma} h_{\gamma\delta;\sigma} - \frac{1}{8} h^{\gamma\delta;\sigma} h_{\gamma\sigma;\delta} + \frac{1}{8} h^{\gamma\gamma;\delta} h^{\delta\sigma}_{;\sigma} \right. \right. \\
 & - \frac{1}{16} h^{\gamma\gamma;\delta} h_{\sigma;\delta}^{\sigma;\delta} \left. \right) + h^\alpha_\alpha h^{\beta\gamma} \left(-\frac{1}{2} h_{\beta\gamma;\delta} h^{\delta\sigma}_{;\sigma} + \frac{1}{2} h_{\beta\gamma;\delta} h_{\sigma;\delta}^{\sigma;\delta} - \frac{1}{2} h^{\delta}_{\delta;\rho} h^\sigma_{\sigma;\gamma} \right. \\
 & + \frac{1}{4} h^{\delta}_{\delta;\rho} h^\sigma_{\sigma;\gamma} + h^{\delta}_{\rho;\sigma} h^\sigma_{\delta;\gamma} - \frac{1}{4} h^{\delta\sigma}_{;\rho} h_{\delta\sigma;\gamma} - \frac{1}{2} h^{\delta}_{\rho;\sigma} h_{\delta\gamma}{}^{\sigma} - \frac{1}{2} h^{\delta}_{\delta;\sigma} h^\sigma_{\rho;\gamma} \\
 & + \frac{1}{2} h_{\rho\delta;\sigma} h_{\gamma}{}^{\sigma;\delta} \left. \right) + h^\alpha_\rho h^{\beta\gamma} \left(h^{\delta}_{\delta;\rho} h^\sigma_{\sigma;\gamma} - h_{\alpha\gamma;\delta} h_{\sigma;\delta}^{\sigma;\delta} + \frac{1}{2} h^{\delta\sigma}_{;\alpha} h_{\delta\sigma;\gamma} \right. \\
 & - h^{\delta}_{\alpha;\sigma} h^\sigma_{\gamma;\delta} - 2h^{\delta}_{\alpha;\sigma} h^\sigma_{\delta;\gamma} + h_{\alpha\gamma;\delta} h^{\delta\sigma}_{;\sigma} + h^{\delta}_{\delta;\alpha} h^\sigma_{\gamma\sigma} - \frac{1}{2} h^{\delta}_{\delta;\alpha} h^\sigma_{\sigma;\gamma} \\
 & + h^{\delta}_{\alpha;\sigma} h_{\gamma\delta}{}^{\sigma} \left. \right) + h^{\alpha\gamma} h^{\beta\delta} \left(h_{\alpha\gamma;\beta} h_{\delta;\sigma}^{\sigma;\delta} - h_{\alpha\gamma;\delta} h_{\sigma;\beta}^{\sigma;\delta} + \frac{1}{2} h_{\alpha\beta;\sigma} h_{\gamma\delta}{}^{\sigma} \right. \\
 & - \frac{1}{2} h_{\alpha\gamma;\sigma} h_{\beta\delta}{}^{\sigma} + h^{\sigma}_{\alpha;\beta} h_{\gamma\sigma;\delta} - h^{\sigma}_{\alpha;\beta} h_{\delta\sigma;\gamma} + h_{\alpha\beta;\delta} h_{\sigma;\gamma}^{\sigma} - 2h^{\sigma}_{\alpha;\beta} h_{\delta\gamma;\sigma} \\
 & + h_{\alpha\gamma;\sigma} h_{\beta;\delta}^{\sigma} \left. \right) + R_{\alpha\beta} \left(-2h^{\alpha\gamma} h_{\gamma\delta} h^{\delta\sigma} h_{\sigma}{}^\beta + h^{\gamma\gamma} h^{\alpha\delta} h_{\delta\sigma} h^{\sigma\beta} + \frac{1}{2} h^{\alpha\gamma} h_{\gamma}{}^\beta h^{\delta\sigma} h_{\delta\sigma} \right. \\
 & - \frac{1}{4} h^{\alpha\gamma} h_{\gamma}{}^\beta h^{\delta}_{\delta} h^{\sigma}_{\sigma} + \frac{1}{3} h^{\alpha\beta} h^{\gamma\delta} h_{\delta\sigma} h^{\sigma}_{\gamma} - \frac{1}{4} h^{\alpha\beta} h^{\gamma\gamma} h^{\delta\sigma} h_{\delta\sigma} + \frac{1}{24} h^{\alpha\beta} h^{\gamma\gamma} h^{\delta}_{\delta} h^{\sigma}_{\sigma} \left. \right) \\
 & + R \left(-\frac{1}{192} h^\alpha_\alpha h^\beta_\beta h^{\gamma\gamma} h^{\delta\delta} + \frac{1}{16} h^\alpha_\alpha h^\beta_\beta h^{\gamma\delta} h_{\gamma\delta} + \frac{1}{4} h^{\alpha\beta} h_{\beta\gamma} h^{\gamma\delta} h_{\delta\alpha} \right. \\
 & \left. - \frac{1}{16} h^{\alpha\beta} h_{\alpha\beta} h^{\gamma\delta} h_{\gamma\delta} - \frac{1}{6} h^\alpha_\alpha h^{\beta\gamma} h_{\gamma\delta} h^{\delta\beta} \right) \left. \right\}
 \end{aligned}$$

Imagine having to do calculations with these interaction vertices!

Perturbative quantization

Want to evaluate the “path integral” for the theory “perturbatively”

Compare to
$$\int_{-\infty}^{+\infty} dx e^{-\frac{1}{2}\alpha x^2 - \frac{1}{4!}\lambda x^4} = \int_{-\infty}^{+\infty} dx \sum_{n=0}^{\infty} \frac{1}{n!} \left(-\frac{1}{4!}\lambda x^4\right)^n e^{-\frac{1}{2}\alpha x^2}$$

$$= \sum_{n=0}^{\infty} \frac{1}{n!} \left(-\frac{1}{4!}\lambda \left(\frac{d}{dy}\right)^4\right)^n \int_{-\infty}^{+\infty} dx e^{-\frac{1}{2}\alpha x^2 + xy} \Big|_{y=0}$$

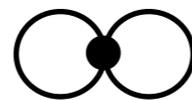
$$= \sqrt{\frac{2\pi}{\alpha}} \sum_{n=0}^{\infty} \frac{1}{n!} \left(-\frac{1}{4!}\lambda \left(\frac{d}{dy}\right)^4\right)^n e^{y^2/2\alpha} \Big|_{y=0}$$

field theory in
zero dimensions
graph combinatorics

On the other hand, setting $\alpha = 1$

$$= \sqrt{\frac{2\pi}{\alpha}} \left(\sum_{\text{4-valent graphs}} \frac{1}{\text{symmetry factor}} \prod_{\text{edges}} \frac{1}{\alpha} \prod_{\text{vertices}} (-\lambda) \right)$$

$$= \sqrt{\frac{3}{\lambda}} e^{\frac{3}{4\lambda}} K_{1/4} \left(\frac{3}{4\lambda}\right)$$



expanding $\frac{1}{\sqrt{2\pi}} \times \text{Integral} = 1 - \frac{\lambda}{8} + \frac{35\lambda^2}{384} - \frac{385\lambda^3}{3072} + \frac{25025\lambda^4}{98304} + O(\lambda^5)$

Perturbative quantization in field theory

Want to compute “correlation functions”

$$\langle h_{\mu_1\nu_1}(x_1) \dots h_{\mu_n\nu_n}(x_n) \rangle := \int \mathcal{D}h h_{\mu_1\nu_1}(x_1) \dots h_{\mu_n\nu_n}(x_n) e^{iS[h]}$$

interpreting this as a Gaussian integral plus perturbation

Similar sum over graphs with “Feynman rules”

$$\frac{1}{\alpha} \rightarrow G_{\mu\nu,\alpha\beta}(x-y) \quad \text{“propagator”}$$

$$\text{satisfies} \quad \left(\delta_{\mu}^{\rho} \delta_{\nu}^{\sigma} - \frac{1}{2} \eta_{\mu\nu} \eta^{\rho\sigma} \right) \square G_{\rho\sigma,\alpha\beta}(x-y) = \delta^4(x-y)$$

Need to multiply propagators, vertex contributions,
then integrate over positions of vertices

Scattering amplitudes

Fourier transformed correlation functions

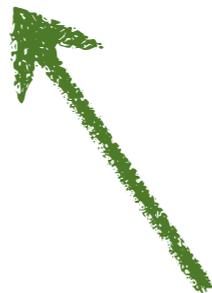
$$\langle h_{\mu_1\nu_1}(k_1) \dots h_{\mu_n\nu_n}(k_n) \rangle$$

Have simple poles at $k_i^2 = 0$

Residues at those poles are graviton scattering amplitudes

Projecting on polarization tensors

$$\mathcal{M}(1^{h_1} \dots n^{h_n}) := \epsilon_{\mu_1\nu_1}^{h_1}(k_1) \dots \epsilon_{\mu_n\nu_n}^{h_n}(k_n) \langle h^{\mu_1\nu_1}(k_1) \dots h^{\mu_n\nu_n}(k_n) \rangle$$



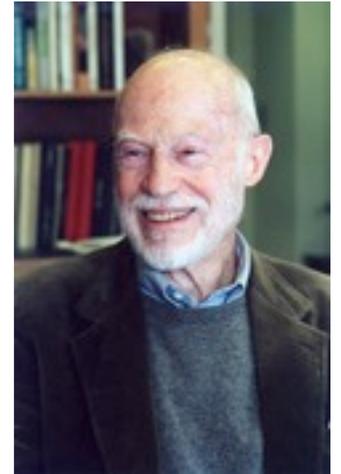
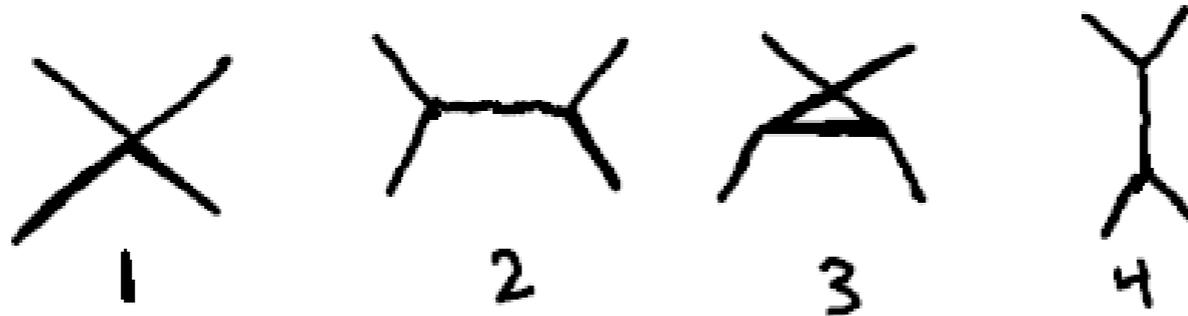
where $h_i = \pm$ “helicity”

$\epsilon_{\mu\nu}^{\pm}(k)$ polarization tensors

Objects of main interest

First calculations

In 1963 I gave [Walter G. Wesley] a student of mine the problem of computing the cross section for a graviton-graviton scattering in tree approximation, for his Ph.D. thesis. The relevant diagrams are these:



Given the fact that the vertex function in diagram 1 contains over 175 terms and that the vertex functions in the remaining diagrams each contain 11 terms, leading to over 500 terms in all, you can see that this was not a trivial calculation, in the days before computers with algebraic manipulation capacities were available. And yet the final results were ridiculously simple.

In modern notations

$$\mathcal{M}(1^-, 2^-, 3^+, 4^+) = \frac{1}{4} s_{12} \frac{s_{12}}{s_{23}} \frac{s_{12}}{s_{24}}$$

where $s_{ij} := (k_i + k_j)^2$

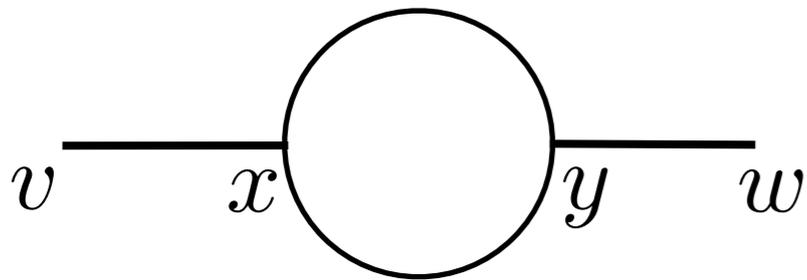
From: Bryce DeWitt
[arXiv:0805.2935](https://arxiv.org/abs/0805.2935)
Quantum Gravity,
Yesterday and Today

Amplitudes for more gravitons are very difficult to obtain
- too many diagrams to consider

Renormalization

In any field theory “loop diagrams” diverge

E.g. scalar field with interaction $\frac{\lambda}{3!}\phi^3$



$$\Delta(v - w) := \frac{1}{2} \int d^4x d^4y P(v - x) P(x - y)^2 P(y - w)$$

Propagator $P(x - y)$ is a distribution

$$\square P(x - y) = \delta^4(x - y)$$

Product of propagators is ill-defined

Need to “regularize”

Dimensional regularization

Gelfand 50's

$$P(x) \sim \frac{1}{|x|^2}$$

Fourier transform

$$\int d^d x \frac{1}{|x|^{2\alpha}} e^{ikx} = \pi^{d/2} \frac{\Gamma(d/2 - \alpha)}{\Gamma(\alpha)} \left(\frac{1}{4} k^2\right)^{\alpha - d/2}$$

This has poles when $\alpha - d/2 = n = 0, 1, \dots$

$$\frac{1}{|x|^{2\alpha}} \sim \frac{1}{d/2 - \alpha + n} \frac{\pi^{d/2}}{\Gamma(d/2 + n)} \frac{1}{2^{2n} n!} \square^n \delta^d(x)$$

E.g. $d = 4 - \epsilon$

$$P^2(x) \sim \frac{1}{\epsilon} \delta^4(x) + \text{non-singular terms}$$

Renormalization

The singular “divergent” part of diagrams is “local”, and can be absorbed into a “renormalization” of parameters (fields)

multiplicative
renormalization

$$\begin{aligned}\phi &\rightarrow Z_\phi \phi \\ \lambda &\rightarrow Z_\lambda \lambda\end{aligned}$$
$$Z = 1 + \frac{c}{\epsilon} + \dots$$

After this is done, finite parts are unambiguously computed in terms of renormalized parameters

Non-trivial mathematical structure: Connes and Kreimer Hopf algebra behind renormalization

This is enough in “renormalizable” field theories

Non-renormalizability of quantum gravity

“Dimensionful” coupling constant $\kappa \sim \sqrt{G} \sim \text{Length}$

Can do more interesting renormalizations

$$h_{\mu\nu} \rightarrow Z_h h_{\mu\nu} + \kappa^2 \left(\underbrace{\frac{\alpha}{\epsilon} R_{\mu\nu}(h) + \frac{\beta}{\epsilon} \eta_{\mu\nu} R(h)}_{\text{new}} \right) + \dots$$

This is sufficient at 1-loop order

‘t Hooft, Veltman '74'

At two loops there is a divergence that cannot be removed by any renormalization

Goroff, Sagnotti '86
van de Ven '91

very difficult calculation:
numerical work

$$\frac{\kappa^2}{\epsilon} \int (W_{\mu\nu}{}^{\rho\sigma})^2$$

$W_{\mu\nu\rho\sigma}$ - Weyl tensor

Need to add this term to the Lagrangian, and then an infinite number of other terms

Part II: New developments

Scattering amplitudes

Witten '03
twistor string

Consider tree-level graviton
scattering amplitude $\mathcal{M}(1^{h_1} \dots n^{h_n})$

Can be seen to be a meromorphic function of k_i with $k_i^2 = 0$

with simple poles at $\left(\sum_{i \in I} k_i\right)^2 = 0$

I - some subset
of momenta

Can show that vanishes when all h_i are plus or all are minus

At least one plus and at least one minus

Label $1^-, 2^+$

Choose

q necessarily complex

$$q : q^2 = 0, \quad k_{1,2} \cdot q = 0$$

Then helicities can be chosen to be

$$\epsilon_{\mu\nu}^-(k_1) = \epsilon_{\mu\nu}^+(k_2) = q_\mu q_\nu$$

BCFW analytic continuation

Britto, Cachazo,
Feng, Witten '05

Continue $k_1 \rightarrow k_1(z) = k_1 + zq$
 $k_2 \rightarrow k_2(z) = k_2 - zq$

Clearly $\sum_i k_i = 0$ still holds

Also
 $k_1(z)^2 = k_2(z)^2 = 0$

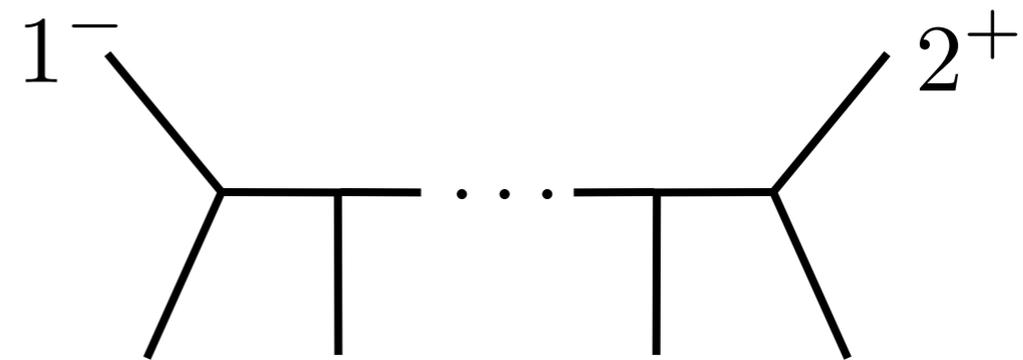
Consider $\mathcal{M}(z)$
Expect

$$\mathcal{M}(z) \sim \frac{z^{2(n-2)}}{z^{(n-3)}} \sim z^{n-1}$$

In fact

$$\mathcal{M}(z) \sim 1/z^2$$

as $z \rightarrow \infty$



maximum $(n - 3)$ propagators in between
 $1/z$ for each propagator
 $(n - 2)$ vertices
 z^2 for each vertex

much “softer” high-energy behaviour than expected

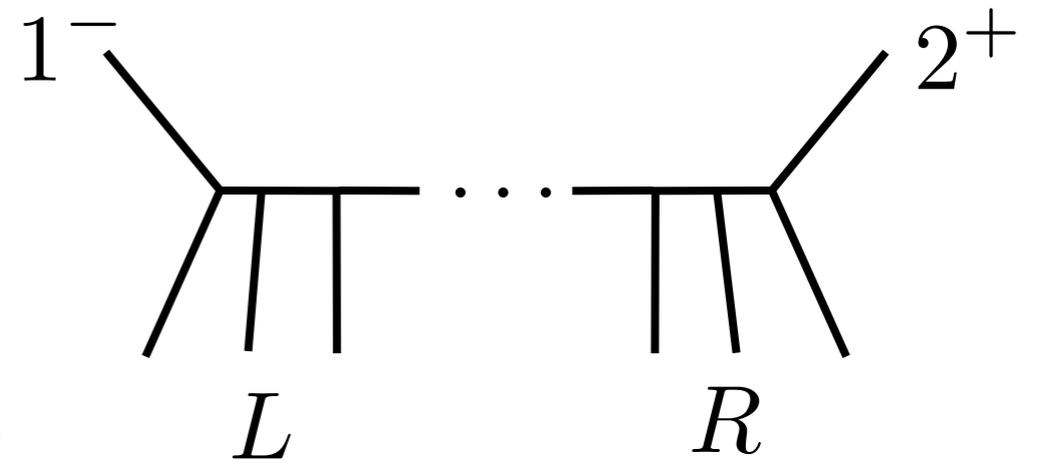
Arkani-Hamed, Kaplan '08

“Softest” UV behaviour known among all QFT’s

Recursion relation

One knows all poles of $\mathcal{M}(z)$

$$0 = \left(k_1 + zq + \sum_{i \in L} k_i \right)^2 = \left(q \cdot \sum_{i \in L} k_i \right) (z - z_L)$$



Residues are

$$\mathcal{R}(z_L) = \frac{1}{\left(q \cdot \sum_{i \in L} k_i \right)} \mathcal{M}_L(z_L) \mathcal{M}_R(z_L)$$

where

$$z_L = - \left(\sum_{i \in 1+L} k_i \right)^2 / \left(q \cdot \sum_{i \in L} k_i \right)$$

Have

$$\int_{|z|=\infty} \frac{dz}{z} \mathcal{M}(z) = 0 = \mathcal{M}(0) + \sum_{z_L} \frac{\mathcal{R}(z_L)}{z_L}$$

amplitudes for smaller number of particles

because $\mathcal{M}(z) \rightarrow 0$ as $z \rightarrow \infty$

$$\mathcal{M}(0) = \sum_L \frac{\mathcal{M}_L(z_L) \mathcal{M}_R(z_L)}{\left(\sum_{i \in 1+L} k_i \right)^2}$$

can get any amplitude recursively from the 3-graviton ones

Explicit formula for any $\mathcal{M}(1^{h_1}, \dots, n^{h_n})$

Cachazo, He, Yuan '13

Consider n equations on an n -punctured sphere

$$\sum_{j \neq i} \frac{s_{ij}}{z_i - z_j} = 0 \quad \text{“scattering equations”}$$

$z_i, i = 1, \dots, n$
puncture locations

$$s_{ij} = (k_i + k_j)^2$$

Using $\sum_i k_i = 0, \quad k_i^2 = 0$ easy to show

- only $(n-3)$ linearly independent
- $\text{SL}(2, \mathbb{C})$ invariant

Can also show that $(n-3)!$ solutions

$$z_{ij} := z_i - z_j$$

Claim

$$\mathcal{M}(\{k, h\}) = \int \frac{\prod_i d^2 z_i}{\text{vol SL}(2, \mathbb{C})} \prod_i \delta' \left(\sum_{j \neq i} \frac{s_{ij}}{z_{ij}} \right) \underline{E^2(\{k, h, z\})}$$

measure

integrand

Measure can be shown to reduce to

$$\sum_{\text{solutions}} \frac{(z_{pq})(z_{qr})(z_{rp})(z_{ij})(z_{jk})(z_{ki})}{|\Phi|_{pqr}^{ijk}}$$

where $|\Phi|_{pqr}^{ijk}$ minor obtained by removing 3 rows 3 columns from

$$\Phi_{ij} = \begin{cases} s_{ij}/z_{ij}^2, & i \neq j \\ -\sum_{k \neq i} s_{ik}/z_{ik}^2, & i = j \end{cases} \quad \begin{array}{l} \text{sum of any row or column zero} \\ \text{compare matrix tree theorems} \end{array}$$

Integrand

$$E^2(\{k, h, z\}) = \frac{1}{z_{ij}^2} |\Psi|_{ij}^{ij}$$

$$\Psi = \begin{pmatrix} A & B \\ -B^T & C \end{pmatrix}$$

anti-symmetric $2n \times 2n$ matrix

$$A_{ij} = \begin{cases} s_{ij}/z_{ij}, & i \neq j \\ 0, & i = j \end{cases} \quad B_{ij} = \begin{cases} (q_i - k_j)^2/z_{ij}, & i \neq j \\ -\sum_{l \neq i} (q_i - k_l)^2/z_{il}, & i = j \end{cases}$$

$$C_{ij} = \begin{cases} (q_i - q_j)^2/z_{ij}, & i \neq j \\ 0, & i = j \end{cases}$$

where q's are "square roots" of polarization tensors $\epsilon^{\mu\nu}(k_i) = q_i^\mu q_j^\nu$

Scattering amplitudes summary

- At large (complex) momenta graviton scattering amplitudes are much better behaved than naive arguments suggest
- Gravity is best behaved QFT in this sense
But it is also worst behaved - non-renormalizability
- Graviton scattering amplitudes can be obtained recursively
BCFW recursion relations
- Closed formula for tree-level amplitudes is possible

This means that one can characterize the space of (perturbative) solutions of GR completely

Space of solutions = phase space

Quantization?

Gauge-theoretic formulation of GR

Given an $SU(2)$ connection A^i
one can define a spacetime metric

connection as a
“potential” for the metric

This metric owes its existence to the isomorphism

$$SO(6, \mathbb{C}) \sim SL(4, \mathbb{C})$$

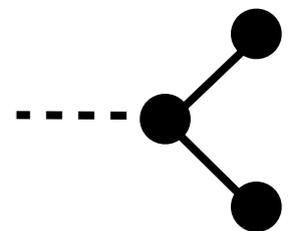


Very important for
twistor theory

Dynkin diagrams



$\mathfrak{sl}(n+1)$



$\mathfrak{so}(2n)$

Proof: Consider the 6-dimensional space Λ^2 of 2-forms in \mathbb{R}^4

The wedge product makes Λ^2 into a metric space

$$\Lambda^2 \ni U, V \rightarrow (U, V) = U \wedge V / d^4x \in \mathbb{R}$$

metric of signature (3,3) if over \mathbb{R}

$SL(4, \mathbb{R})$ acts on Λ^2 $G_{\mu}^{\nu} \in SL(4, \mathbb{R})$

$${}^G U_{\mu\nu} = G_{\mu}^{\alpha} G_{\nu}^{\beta} U_{\alpha\beta}$$

the wedge product metric is preserved

$$\Rightarrow SL(4, \mathbb{R}) \sim SO(3, 3)$$

The isomorphism implies

$$SL(4)/SO(4)$$

$$SO(3,3)/SO(3) \times SO(3)$$



conformal
metrics on M

Grassmanian of
3-planes in Λ^2

Conformal metrics can be encoded into the
knowledge of which 2-forms are self-dual

Explicitly: a triple of linearly independent 2-forms $B_{\mu\nu}^i$

$$\Rightarrow g_{\mu\nu} \sim \tilde{\epsilon}^{\alpha\beta\gamma\delta} \epsilon^{ijk} B_{\mu\alpha}^i B_{\nu\beta}^j B_{\gamma\delta}^k$$

Urbantke
formula

2-forms $B_{\mu\nu}^i$ are self-dual with respect to this metric

Definition of the metric:

Let A^i be an $SU(2)$ connection

$$F^i = dA^i + (1/2)[A, A]^i$$

$\left(\begin{array}{l} SL(2, \mathbb{C}) \text{ connection for} \\ \text{Lorentzian signature} \end{array} \right)$

$$F^i \wedge (F^j)^* = 0$$

reality conditions

declare F^i to be self-dual 2-forms \Rightarrow conformal metric

To complete the definition of
the metric need to specify
the volume form

$$(\text{vol}) := \frac{1}{\Lambda^2} f(F \wedge F)$$

$$\Lambda \sim 1/L^2$$

dimensionful parameter

Functions of the curvature

Let f be a function on $\mathfrak{g} \otimes_S \mathfrak{g}$ satisfying

\mathfrak{g} - Lie algebra of G

$f : X \rightarrow \mathbb{R}(\mathbb{C})$ defining function

$X \in \mathfrak{g} \otimes_S \mathfrak{g}$

1) $f(\alpha X) = \alpha f(X)$

homogeneous degree 1

2) $f(\text{Ad}_g X) = f(X), \quad \forall g \in G$

gauge-invariant

Then $f(F \wedge F)$ is a well-defined 4-form (gauge-invariant)

Choose a volume form and define X^{ij}

$$F^i \wedge F^j := X^{ij}(\text{vol})$$

then $f(F \wedge F) := (\text{vol}) f(X)$

independent of choice of (vol)

To motivate a choice of $f(X)$

take an Einstein metric, consider

the Levi-Civita A^i connection on $\Lambda^+ \subset \Lambda^2$

the space of self-dual 2-forms

$$F^i = \left(\frac{\Lambda}{3} + W^+ \right)^{ij} \Sigma^j \quad \text{Tr}(W^+) = 0$$

where Σ^i is a basis of self-dual 2-forms

Then

$$\left(\text{Tr} \sqrt{F \wedge F} \right)^2 = 2\Lambda^2(\text{vol}) \quad \Sigma^i \wedge \Sigma^j \sim \delta^{ij}$$

This suggest that we take

$$(\text{vol}) := \frac{1}{2\Lambda^2} \left(\text{Tr} \sqrt{F \wedge F} \right)^2$$

this completes the definition of the metric from A^i

Variational principle

Consider a functional that is just a multiple of the volume

$$S[A] = \frac{\Lambda}{8\pi G} \int (\text{vol})$$

$\Lambda \neq 0$ KK PRL106:251103,2011

related ideas for zero scalar
curvature in early 90's
Capovilla, Dell, Jacobson

Critical points

$$(*) \quad d_A \left(\text{Tr} \sqrt{X} (X^{-1/2})^{ij} F^j \right) = 0$$

second-order PDE's for the connection

Theorem:

For connections A^i satisfying (*)

the metric $g(A)$ is Einstein with non-zero scalar curvature Λ

In the opposite direction, the self-dual part of the Levi-Civita connection for an Einstein metric satisfies (*)

Caveat: only metrics with $\Lambda/3 + W^+$
invertible almost everywhere covered

examples not
covered
 $S^2 \times S^2$
Kahler metrics

Gauge-theoretic perturbation theory

Need a non-zero connection to expand about
Homogeneous isotropic connection

$$A^i = i a(t) dx^i$$

$$F^i \wedge F^j \sim \delta^{ij}$$

such a connection is a solution

The corresponding metric is de Sitter of cosmological constant Λ
(in flat slicing)

$$F^i = \frac{\Lambda}{3} \Sigma^i$$

where Σ^i is a basis of self-dual 2-forms

One gets the following **linearised Lagrangian**

square of a certain Dirac operator

$$L^{(2)} \sim P_{ijkl}^{(2)} (\Sigma^{i\mu\nu} d_\mu a_\nu^j) (\Sigma^{k\rho\sigma} d_\rho a_\sigma^l)$$

where

$$P_{ijkl}^{(2)} := \delta_{i(k} \delta_{l)j} - \frac{1}{3} \delta_{ij} \delta_{kl}$$

a_μ^i connection perturbation

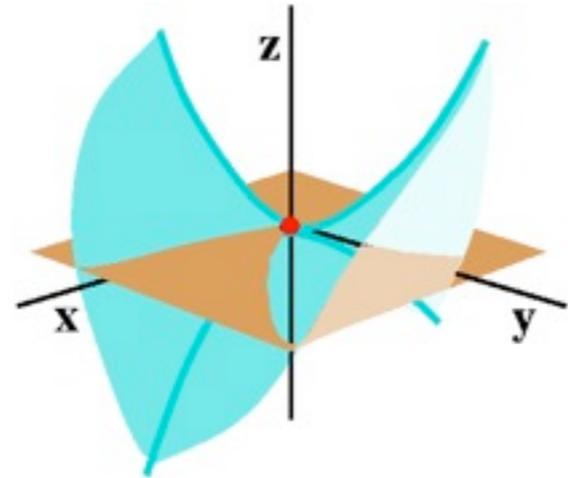
d_μ de Sitter covariant derivative

Easy to show that describes gravitons on de Sitter space
considerably simpler linearization than in the metric case

The new formulation is simpler than the metric-based GR

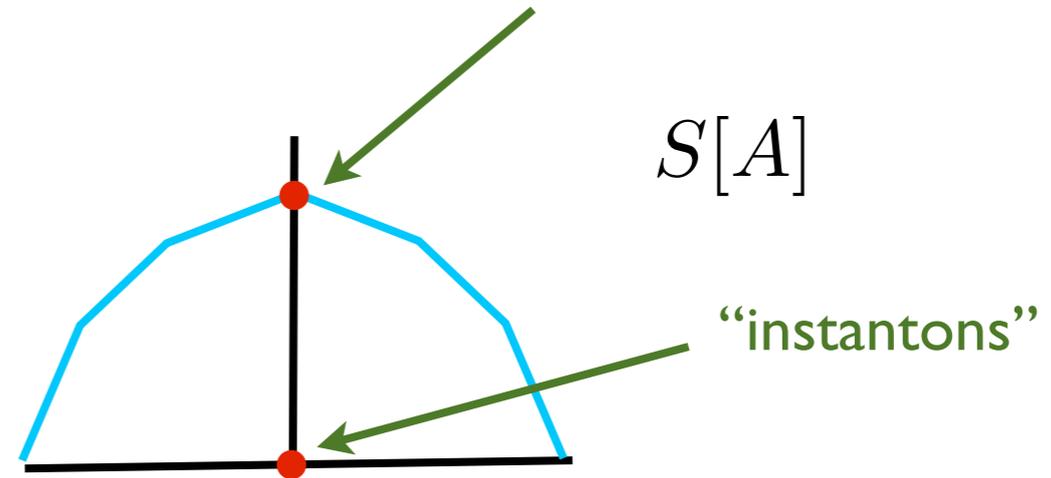
Concave action functional

$S_{\text{EH}}[g]$



space of metrics

$2\tau(M) + 3\chi(M)$



SU(2) connections/gauge

α, β - dimensionless parameters

$$\mathcal{L}^{(3)} = \frac{\alpha}{M^2} \text{Tr} \left((\Sigma da)^3 \right) + \frac{\beta}{M^2} (\Sigma da)^{ij} \left(\frac{1}{i} \epsilon^{\mu\nu\rho\sigma} d_\mu a_\nu^i d_\rho a_\sigma^j + M^2 \Sigma^{i\mu\nu} \epsilon^{jkl} a_\mu^j a_\nu^k \right)$$

with $(\Sigma da)_{ij} := P_{ijkl}^{(2)} (\Sigma^{k\mu\nu} d_\mu a_\nu^l)$ $M^2 := \Lambda/3$

compare with the mess in the metric formulation

Summary:

- After a long period of inactivity, perturbative quantum gravity is again at the cross-roads of many interesting developments
- Gravity has much better high-energy behaviour than was thought
- Powerful recursion relations for the amplitudes
- Tree-level amplitudes can be solved for in closed form
- GR can be described as an $SU(2)$ gauge theory of a novel type:
bounded from above Euclidean action
- Produces much simpler perturbative expansion than
the usual description

Further interesting developments are guaranteed

Thank you!

Quantum Theory Hopes

Remark: no dimensionful coupling constants
in any of these gravitational theories

(negative) dimension coupling
constant comes when expanded
around a background



Non-renormalizable in the usual sense

Hope: the class of theories {all possible $f()$ } is large enough
to be closed under renormalization

$$\frac{\partial f(F \wedge F)}{\partial \log \mu} = \beta_f(F \wedge F)$$

I.e. physics at higher energies continues to be
described by theories from the same family

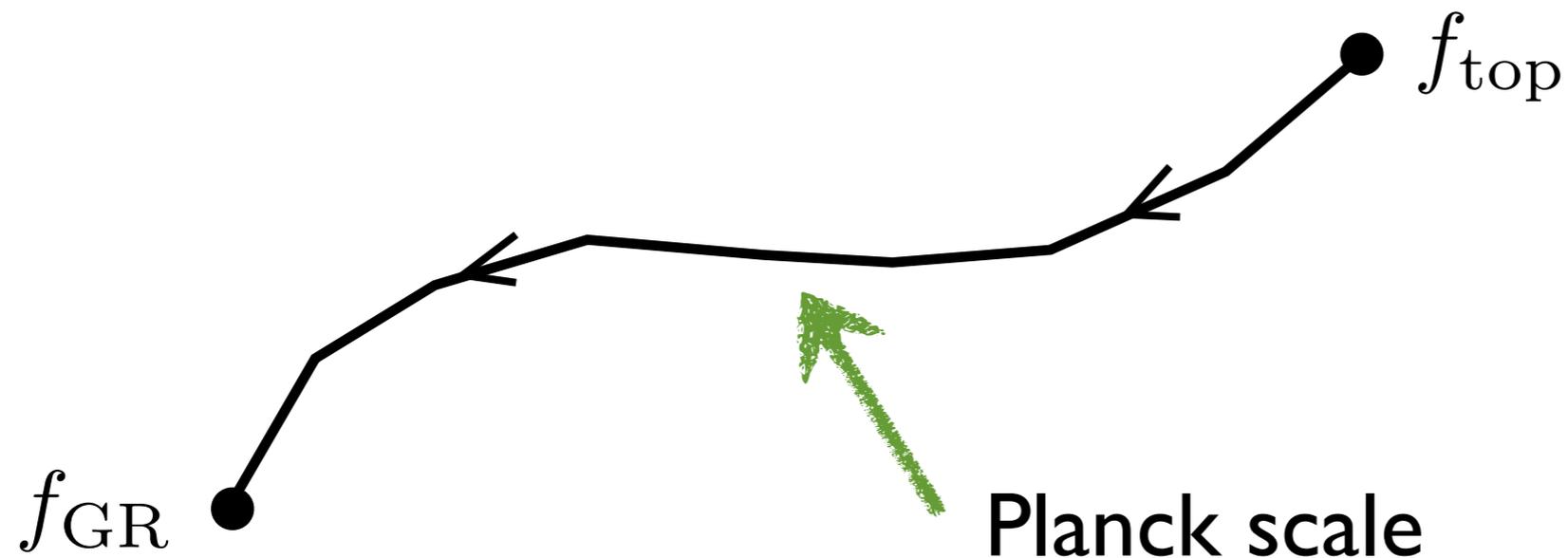
= no new DOF appear
at Planck scale, just the
dynamics changes

The speculative RG flow: topological theory ?

$$f_{\text{top}}(F \wedge F) = \text{Tr}(F \wedge F)$$

necessarily a fixed point
of the RG flow

corresponds to a topological theory
(no propagating DOF)



flow from very steep
in IR towards very
flat in UV potential